

Chandler Wobble and Free Core Nutation: Theory and Features

Sung-Ho Na^{1,2†}, Kyoung-Min Roh¹, Jungho Cho¹, Sung-Moon Yoo¹, Byungkyu Choi¹,
 Hasu Yoon¹

¹Korea Astronomy and Space Science Institute, 776 Daedeok-daero, Yuseong-gu, Daejeon, 34055, Korea

²University of Science and Technology, 217 Gajung-ro, Yuseong-gu, Daejeon, 34113, Korea

Being a torque free motion of the rotating Earth, Chandler wobble is the major component in the Earth's polar motion with amplitude about 0.05-0.2 arcsec and period about 430-435 days. Free core nutation, also called nearly diurnal free wobble, exists due to the elliptical core-mantle boundary in the Earth and takes almost the whole part of un-modelled variation of the Earth's pole in the celestial sphere beside precession and nutation. We hereby present a brief summary of their theories and report their recent features acquired from updated datasets (EOP C04 and ECMWF) by using Fourier transform, modelling, and wavelet analysis. Our new findings include (1) period-instability of free core nutation between 420 and 450 days as well as its large amplitude-variation, (2) re-determined Chandler period and its quality factor, (3) fast decrease in Chandler amplitude after 2010.

Keywords: Chandler wobble, free core nutation

1. INTRODUCTION

Chandler wobble and free core nutation are two major modes of perturbation in the Earth rotation. Earth rotation status needs to be known for the coordinate conversion between celestial reference frame and terrestrial reference frame. Due mainly to the tidal torque exerted by the moon and the sun on the Earth's equatorial bulge, the Earth undergoes precession and nutation. Modeling of precession and nutation has become rigorous ever since the first development of Woolard and later one of Kinoshita (Moritz & Mueller 1988). The Earth's reference pole, which is the origin of the geographic coordinates; latitude and longitude, was determined in the beginning of 20th century. However, the Earth's rotational pole moves in time on the surface of the Earth, and this motion cannot be easily modelled because of nonlinear behavior of atmosphere and other various movements in the Earth. So called polar motion refers to the position of the true spin rotational pole denoted as (xp, yp). Official observation of the polar motion started more than a century years ago. Chandler wobble

is also called Eulerian free nutation, because Euler early predicted its existence. Chandler wobble, along with annual wobble, is the largest component of the polar motion, and its period is about 430-435 days. Theoretical investigations and observations on its characters - period, quality factor, and its excitation source have been made repeatedly with improvements (Gross 2009). Poincot diagram of Chandler wobble and the locus of angular velocity vector on the Earth during one Chandler period are illustrated in Fig. 1 (left side). The body cone rotates around space cone every day. On the Earth's surface the pole of rotation slowly moves counterclockwise around the symmetry axis of the Earth with the Chandler period - here annual wobble or others are not considered. Smith (1977) clearly confirmed the definition of wobble and nutation: 'wobble' is the angle between CIP and symmetry axis, and 'nutation' is the angle of the symmetry axis from angular momentum vector. This definition was first made by Munk & MacDonald (1960).

Not only the pole position changes on the Earth's surface, but also exist small perturbation in the pole position on the celestial sphere - imposed on the known precession and

© This is an Open Access article distributed under the terms of the Creative Commons Attribution Non-Commercial License (<https://creativecommons.org/licenses/by-nc/3.0/>) which permits unrestricted non-commercial use, distribution, and reproduction in any medium, provided the original work is properly cited.

Received 21 JAN 2019 Revised 4 MAR 2019 Accepted 8 MAR 2019

†Corresponding Author

Na SH is now working as lecturer at universities.

Tel: +82-70-4284-0702, E-mail: sunghona@kasi.re.kr

ORCID: <https://orcid.org/0000-0003-2004-0715>

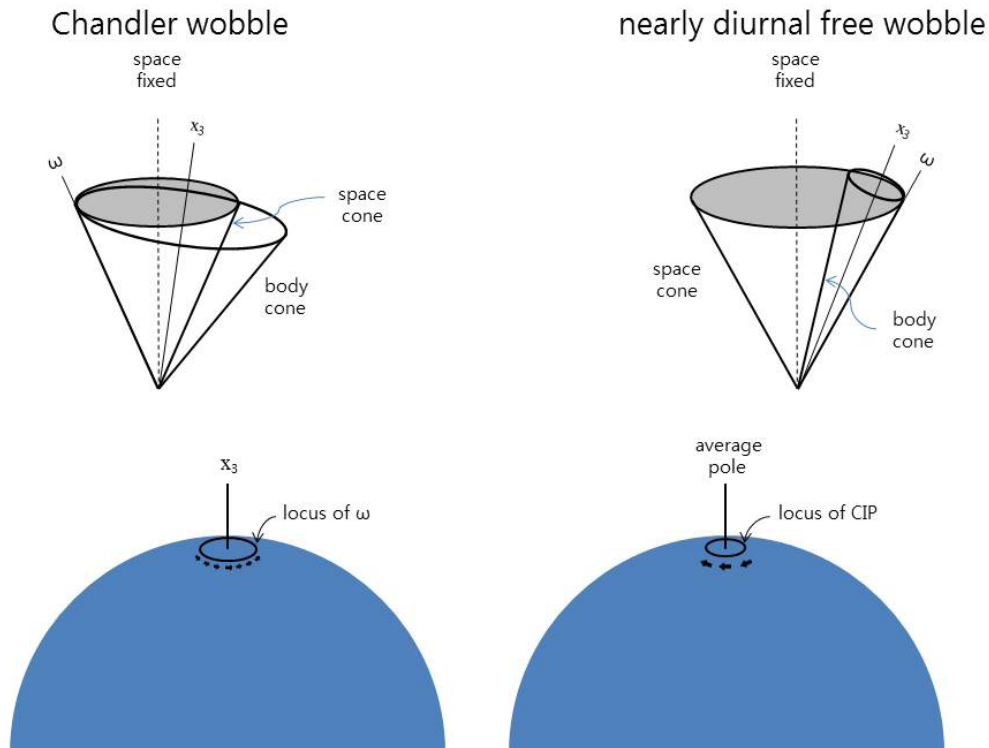


Fig. 1. Schematic illustrations of Chandler wobble and free core nutation (nearly diurnal free wobble). Upper ones are Poincaré diagrams, while lower ones are pole traces on the Earth. Daily rotation is described by the rotating motion of body cone around space cone. Chandler wobble is a prograde motion around symmetry axis of the Earth (period about 435 days). Nearly diurnal free wobble is the retrograde daily motion of the pole on the Earth, while the Earth's symmetry axis slowly rotates clockwise with maintaining the angles unchanged – the latter is free core nutation. For both cases, wobble is the angle between CIP and the Earth's symmetry axis, and nutation is the angle of the symmetry axis from angular momentum vector (space fixed).

nutations. This un-modelled error has been suspected to be associated with the free core nutation, which is another intrinsic mode of Earth's rotational perturbation due to the mismatch between the two each rotational axes of the Earth's mantle and core. After successful VLBI operation as well as accurate measurement of tidal gravity since 1980s, free core nutation has become clearly observable. The period of free core nutation has been reported to be around 430 days in the celestial sphere, however, free core nutation appears to be a diurnal variation for the observers on the Earth. Therefore it is also called nearly diurnal free wobble. Poincaré diagram of free core nutation and its diurnal locus of angular velocity vector on the Earth are illustrated in Fig. 1 (right side). The body cone spins clockwise nearly daily and rotates clockwise slowly with contacting the space cone: nearly diurnal free wobble. In this case the nutation angle is much larger than wobble (400 – 460 times) and body cone slowly moves with contacting the space cone with much longer period (again 400 – 460 times): free core nutation.

The IERS recommendation of the coordinate conversion from Celestial Reference Frame to Terrestrial Reference Frame is given as follows (Petit & Luzum 2010),

$$\begin{bmatrix} x_1 \\ x_2 \\ x_3 \end{bmatrix}_{\text{Terrestrial}} = R_{pm} R_{spin} R_{prec+nut} \begin{bmatrix} x_1 \\ x_2 \\ x_3 \end{bmatrix}_{\text{Celestial}}$$

where $R_{prec+nut}$, R_{spin} , and R_{pm} are defined as follows.

$$R_{prec+nut} = \begin{pmatrix} 1 - bX^2 & -bXY & -X \\ -bXY & 1 - bY^2 & -Y \\ X & Y & 1 - b(X^2 + Y^2) \end{pmatrix}$$

$$R_{spin} = R_3(\theta - s + s')$$

$$R_{pm} = R_1(-y_p)R_2(-x_p)R_3\left(\frac{x_p y_p}{2}\right)$$

where; (1) X and Y are coordinates of the Earth pole in the celestial sphere - representing precession and nutation, (2) θ , s , and s' are Earth's spin rotation angle and two small correction terms, (3) x_p and y_p are called polar motion and denote true pole position on the Earth's surface. More description of this coordinate transformation is given in the appendix.

Being the main components in pole positional perturbation on the Earth and celestial sphere,

investigations have been carried on Chandler wobble and free core nutation. One good summary about Chandler wobble was given by Gross (2009). An earlier workshop on Chandler wobble held in 2004 was a special and fruitful occasion dealing with diverse related disciplines (Plag et al. 2005). Mechanical existence of Chandler wobble mode itself is due to the equatorial bulge of the Earth, as early studied by Euler. Its reality was later confirmed by Chandler, while the period discrepancy of 14 months from 305 day was interpreted as the influence of mantle elasticity by Newcomb. Dahlen's theoretical work was an important advance to explain the role of the global oceans increasing the period of Chandler wobble by 28 days. There have been controversies on the excitation mechanism of Chandler wobble, however, as accurate datasets accumulate on the global atmosphere and ocean, it is now widely accepted that Chandler wobble is being excited mainly by the changes in the states of atmosphere and ocean. Water and glacial mass redistribution on the surface and shallow crust of the Earth do another role in the polar motion. Earthquakes are no longer regarded as appreciable energy source of Chandler wobble (Gross 2009; Chung & Na 2016). Accurate estimation of Chandler wobble period and quality factor has long been attempted repeatedly by investigators using different approaches. Statistical methods were employed early, but after global data coverages, comparison between observation and model has become feasible. Furuya & Chao (1996) estimated Chandler period and quality factor as 433.7 days and 49 by minimizing the residual power between geodetic excitation function and atmospheric excitation. Gross elaborated and attained 431.9 days and 83 (Plag et al. 2005). There have been other type approaches, which led more or less similar results (see Plag et al. 2005; Gross 2009).

Theoretical studies of free core nutation started long ago by Poincare, Hough and others (Moritz & Mueller 1988), and later important step was made by Sasao (1980). The torque due to misalignment between the rotational axes of the mantle and the liquid core is known to be the cause of free core nutation. Unlike Chandler wobble excitation, the formulation of free core nutation excitation function is not clearly known. However, it is believed that free core nutation is also excited by atmosphere and ocean. Observation of free core nutation has been made either by accurate gravity observation or VLBI data analysis. Unlike its predicted period as 460 days by Sasao, studies repeatedly confirmed free core nutation period at around 430 days with small deviations (for example, Rajner & Brzezinsky (2017), Krasna et al. (2013), Defraigne et al. (1994) and dozens others cited in their references). On the contrary, large variation of the free core nutation period has been reported by Cummins

& Wahr (1993) and recently by Gubanov (2010). Chao and Hsieh recently reported its period as 445 days (Chao & Hsieh 2015). Reported value of nearly diurnal free wobble Q varied largely from 1000 to several thousands. It is noted here that quality factor of free core nutation and that of nearly diurnal free wobble should be discerned, even though the two are identical phenomenon viewed from different reference frames: $Q_{fcn} T_{fcn} = Q_{ndfw} T_{ndfw}$.

Deep Earth inside processes are also candidates of some parts of Earth rotation variation - particularly decadal and other long periodic variations. All the dynamic processes of the Earth - atmospheric/oceanic/deep inside - eventually affect both Chandler wobble and free core nutation as well as other type perturbations in different amounts. In this article, we briefly summarize the theories of Chandler wobble and free core nutation. Also we report a few interesting results of our spectral analysis and modelling on them based on recent datasets.

2. BRIEF THEORETICAL SUMMARY

Except glacial isostatic adjustment and other slow Earth deep interior processes, overall mechanical response of the rotating Earth in the range of Chandler wobble and free core nutation can be explained by using linearized equation. This linearization holds true, even though the equations of hydrodynamic motions in the atmosphere or ocean are inherently nonlinear. In this section we summarize the theory of perturbation mode in the Earth's spin rotation, which was originally given by Moritz & Mueller (1988).

Starting with rigid Earth we have simple equation for perturbation in its spin rotation as follows.

$$\begin{aligned} A \frac{d\omega_1}{dt} + \omega_2 \omega_3 (C - A) &= 0 \\ A \frac{d\omega_2}{dt} - \omega_3 \omega_1 (C - A) &= 0 \\ C \frac{d\omega_3}{dt} &= 0 \end{aligned} \quad (1)$$

Define a coordinate frame $\vec{r}_0 = (x_0, y_0, z_0)$ rotating with constant angular velocity $\vec{\omega}_0 = (0, 0, \omega_0)$ and another coordinate frame $\vec{r} = (x, y, z)$ which has a slight deflection $\vec{\theta}_0 = (\theta_1, \theta_2, \theta_3)$. Then the relation between the two coordinate frames can be written as $\vec{r} = \vec{r}_0 - \vec{\theta}_0 \times \vec{r}_0$, and the angular velocity of in \vec{r} frame can be written as

$$\begin{bmatrix} \omega_1 \\ \omega_2 \\ \omega_3 \end{bmatrix} = \begin{bmatrix} 0 \\ 0 \\ \omega_0 \end{bmatrix} + \begin{bmatrix} \dot{\theta}_1 \\ \dot{\theta}_2 \\ \dot{\theta}_3 \end{bmatrix} + \begin{bmatrix} -\theta_2 \omega_0 \\ \theta_1 \omega_0 \\ 0 \end{bmatrix} \quad (2)$$

Insert Eq.(2) into Eq.(1), then we find (with setting $\dot{\theta} = 0$),

$$\begin{aligned} A(\ddot{\theta}_1 - \dot{\theta}_2\omega_0) + (C - A)(\dot{\theta}_2 + \theta_1\omega_0)\omega_0 &= 0 \\ A(\ddot{\theta}_2 + \dot{\theta}_1\omega_0) + (A - C)(\dot{\theta}_1 - \theta_2\omega_0)\omega_0 &= 0 \\ \ddot{\theta}_3 &= 0 \end{aligned} \tag{3}$$

After rewriting the first two of Eq.(3) with defining $\Omega_r = \frac{C-A}{A}\omega_0$ and complex angle $\theta = \theta_1 + i\theta_2$, we find

$$\ddot{\theta} + i(\omega_0 - \Omega_r)\dot{\theta} + \Omega_r\omega_0\theta = 0 \tag{4}$$

Assuming the solution for the Eq. (4) as $\theta \propto e^{i\Omega t}$, then we have the resulting condition for Ω as $\Omega^2 + (\omega_0 - \Omega_r)\Omega - r_0 = 0$. Two solution of Ω are found as $\Omega = \Omega_r = \frac{C-A}{A}\omega_0$ and $\Omega = -\omega_0$. These two may be compared with the frequencies of Chandler wobble and retrograde diurnal free wobble.

Now we assume the Earth is composed of rigid mantle and liquid core. Also we assume the core-mantle boundary is slightly elliptical so that it is expressed as follows.

$$\frac{x^2 + y^2}{a^2} + \frac{z^2}{c^2} = 1$$

The dynamical ellipticity of the core is almost the same as its flattening: $\varepsilon = \frac{C_c - A_c}{C_c} = \frac{2a^2 - (a^2 + c^2)}{2a^2} \cong \frac{a-c}{a}$. Denote the perturbed angular velocity of the mantle as $\vec{\omega} = (\omega_1, \omega_2, \omega_3)$ and the relative angular velocity of the core w.r.t. mantle as $\vec{x} = (x_1, x_2, x_3)$, then the kinetic energy associated with the perturbed motion can be written as

$$T = \frac{1}{2}A(\omega_1^2 + \omega_2^2) + \frac{1}{2}C\omega_3^2 + \frac{1}{2}A_c(x_1^2 + x_2^2) + \frac{1}{2}C_c x_3^2 + F(\omega_1 x_1 + \omega_2 x_2) + C_c \omega_3 x_3$$

where $A_c = \frac{1}{5}M_c(a^2 + c^2)$, $C_c = \frac{2}{5}M_c a^2$, and $F = \frac{2}{5}M_c a c$ are the rotational inertia of the core.

Consider torque free motion for the whole Earth, then we have

$$\begin{aligned} \frac{\partial}{\partial t}(A\omega_1 + Fx_1) + \omega_2(C\omega_3 + C_c x_3) - \omega_3(A\omega_2 + Fx_2) &= 0 \\ \frac{\partial}{\partial t}(A\omega_2 + Fx_2) + \omega_3(A\omega_1 + Fx_1) - \omega_1(C\omega_3 + C_c x_3) &= 0 \\ \frac{\partial}{\partial t}(C\omega_3 + C_c x_3) + (x_2\omega_1 - x_1\omega_2)F &= 0 \end{aligned}$$

Likewise we have following equation for the core free motion as

$$\begin{aligned} \frac{\partial}{\partial t}(A_c x_1 + F\omega_1) + x_3(A_c x_2 + F\omega_2) - x_2 C_c(x_3 + \omega_3) &= 0 \\ \frac{\partial}{\partial t}(A_c x_2 + F\omega_2) - x_3(A_c x_1 + F\omega_1) + x_1 C_c(x_3 + \omega_3) &= 0 \\ \frac{\partial}{\partial t}(C_c(x_3 + \omega_3)) + (x_2\omega_1 - x_1\omega_2)F &= 0 \end{aligned}$$

From the third equations of the two sets, we have $\frac{\partial}{\partial t}[(C - C_c)\omega_3] = 0$. Assuming $\omega_1, \omega_2, x_1, x_2$ are small enough, the four other equations can be rewritten as follows.

$$\begin{aligned} A\dot{\omega}_1 + F\dot{x}_1 + (C - A)\omega_0\omega_2 - F\omega_0x_2 &= 0, \\ A\dot{\omega}_2 + F\dot{x}_2 - (C - A)\omega_0\omega_1 + F\omega_0x_1 &= 0 \end{aligned}$$

$$\begin{aligned} A_c\dot{x}_1 + F\dot{\omega}_1 - x_2 C_c\omega_0 &= 0, \\ A_c\dot{x}_2 + F\dot{\omega}_2 + x_1 C_c\omega_0 &= 0 \end{aligned}$$

These can be gathered into two following complex equations with defining complex quantities as $\omega = \omega_1 + i\omega_2$ and $x = x_1 + ix_2$.

$$\begin{aligned} A\dot{\omega} + F\dot{x} - i(C - A)\omega_0\omega + iF\omega_0x &= 0 \\ A_c\dot{x} + F\dot{\omega} + iC_c\omega_0x &= 0 \end{aligned} \tag{5}$$

Assuming oscillatory solution for ω and x ($\propto e^{i\Omega t}$), Eq.(5) is expressed in matrix form as

$$\begin{bmatrix} A\Omega - (C - A)\omega_0 & F(\Omega + \omega_0) \\ F\Omega & A_c\Omega + C_c\omega_0 \end{bmatrix} \begin{bmatrix} \omega \\ x \end{bmatrix} = \begin{bmatrix} 0 \\ 0 \end{bmatrix}$$

Then, for existence of nontrivial solution, we have following condition.

$$(A\Omega - (C - A)\omega_0)(A_c\Omega + C_c\omega_0) - F^2\Omega(\Omega + \omega_0) = 0 \tag{6}$$

With approximations as $C_c \cong A_c(1 + \varepsilon)$ and $F^2 = C_c(2A_c - C_c) \cong A_c^2$, Eq.(6) can be written as

$$(A\Omega - (C - A)\omega_0)(\Omega + (1 + \varepsilon)\omega_0) - A_c\Omega(\Omega + \omega_0) = 0 \tag{7}$$

The frequency Ω of the mode of our interest can be found by using perturbation scheme, for the ellipticity ε is small (less than 1 percent). Taking $\varepsilon = 0$ in Eq. (7), we find two solutions of zero order as follows.

$$\Omega_1^{(0)} = \frac{C - A}{A_m}\omega_0, \quad \Omega_2^{(0)} = -\omega_0 \tag{8}$$

By putting terms of ε into right hand side, Eq. (7) can be rewritten as follows.

$$(A_m\Omega - (C - A)\omega_0)(\Omega + \omega_0) = -(A\Omega - (C - A)\omega_0)\omega_0\varepsilon \tag{9}$$

Write $\Omega_1^{(1)} = \Omega_1^{(0)} + \alpha\varepsilon$ and insert it into Eq. (9), we find $\alpha = \frac{A_c A (C - A)}{A_m^2 C} \omega_0 \cong -0.00047\omega_0$. Corresponding frequency is $\Omega_1^{(1)} = \Omega_1^{(0)} + \alpha\varepsilon = \frac{C - A}{A_m} \omega_0 - 0.00047\varepsilon\omega_0 \cong \Omega_1^{(0)}$. Likewise, write $\Omega_2^{(1)} = \Omega_2^{(0)} + \alpha\varepsilon$ and insert it into Eq. (9), we find $\alpha = -\frac{C}{C - A + A_m} \omega_0 \cong -\frac{C}{C_m} \omega_0$ and so $\Omega_2^{(1)} = \Omega_2^{(0)} + \alpha\varepsilon = -(1 + \varepsilon\frac{C}{C_m})\omega_0$. With proper values of $\varepsilon = 1/400$ and $C/C_m = 1.13$, we find $\varepsilon C/C_m = 0.0028$.

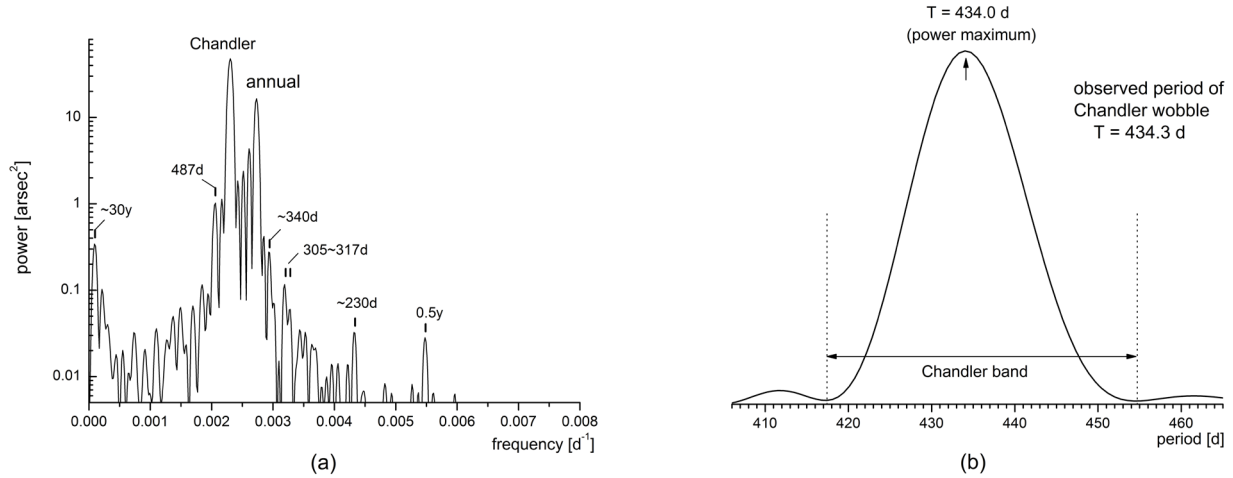


Fig. 2. Polar motion power spectra: (a) Fourier power spectrum of EOP C04 (Jan 1962 – Aug 2018) shown on frequency axis, (b) Same power spectrum closely shown on period axis (Chandler band only). Average Chandler period is found 434.3 days.

$$\Omega_1^{(1)} \cong \Omega_1^{(0)} = \frac{C-A}{A_m} \omega_0, \Omega_2^{(1)} = -(1 + \varepsilon \frac{C}{C_m}) \omega_0 \cong -1.0028 \omega_0 \quad (10)$$

Sasao (1980) extended this theory by including elasticity of mantle. His final equation for the eigenfrequency of perturbation mode of Earth rotation is given as follows.

$$A[\Omega - e\omega_0 + (\Omega + \omega_0)\kappa][(1 + \beta)\Omega + (1 + \varepsilon)\omega_0] - (A_c + \varepsilon A)(1 + \gamma)\Omega(\Omega + \omega_0) = 0$$

Corresponding solutions of Ω are found as follows.

$$\Omega_1 = \frac{A}{A_m}(e - \kappa)\omega_0, \Omega_2 = -(1 + (\varepsilon - \beta)\frac{A}{A_m})\omega_0 \quad (11)$$

The values of parameters used above are known as $e = 0.00329$, $\varepsilon = 0.00252$, $\kappa = 0.00105$, $\beta = 0.00063$. Therefore, the frequencies are $\Omega_1 = 0.00253\omega_0$ and $\Omega_2 = -1.00213\omega_0$. The period of the wobble corresponding to $\Omega_1 = 0.00253\omega_0$ is 396 days, however, the effect of ocean is to increase this value by 28 days. Then theoretical value of Chandler period becomes 424 days. Also the frequency of second mode (nearly diurnal free wobble) is $\Omega_2 = -1.00213\omega_0$, while its apparent frequency to the inertial observer in space (free core nutation) should be $-1.00213\omega_0$, of which period is about 470 days.

3. DATA ANALYSIS

For inspection of recent features of Chandler wobble and free core nutation, we used IERS EOP C04 time series (January 1962 - August 2018), which contain the dataset of polar motion and LOD as well as the celestial pole offset; δX and δY . We evaluated atmospheric excitation function by using the global ECMWF interim dataset of barometric pressure and wind distribution.

3.1 FEATURES OF CHANDLER WOBBLE

Polar motion is referred to the (x_p, y_p) , which is the coordinate of Earth's spin angular velocity (termed as CIP) with respect to its nominal position of 1900-1905. The Fourier power spectrum of the polar motion is shown in Fig. 2. To enhance the resolution and fidelity of the spectrum we separated the daily-basis dataset into four subsets of day number $n = 4k, 4k+1, 4k+2,$ and $4k+3$, and then added all the each four corresponding spectra. Obviously Chandler and annual wobble take the dominant parts in the spectrum, and there also exist other components of which periods are about 487, 340, about 300 days and others (Fig. 2a). Period of maximum power in the Chandler band was found as 434.0 days (Fig. 2b). From analysis on power spectrum in frequency domain, its center period was found as 434.3 days ($423.6d < T < 445.6d$ as $\pm 1\sigma$ range). Assuming white spectral excitation, one may deduce the quality factor from the spectrum. We found corresponding value as $Q = 30$. Chandler wobble components acquired by filtering in the frequency domain from IERS EOP C04 data are shown in Figs. 3a and 3b. It is noticeable that the amplitude of Chandler wobble has been much reduced since 2010. After 2015, its amplitude is less than 30 milliarcsec. Wavelet spectrum of Chandler wobble time series has been attained by using Morlet wavelet and is illustrated in Fig. 4. In the figure, a slight variation of temporal period of Chandler wobbling motion can be seen. 3-dimensional surface illustration wavelet spectrum describes evident Chandler wobble amplitude variation together with its slight periodicity variation.

By comparing geodetic excitation function derived from polar motion and atmospheric excitation function derived from ECMWF dataset in frequency domain of Chandler band, we acquired estimate of Chandler period and quality factor.

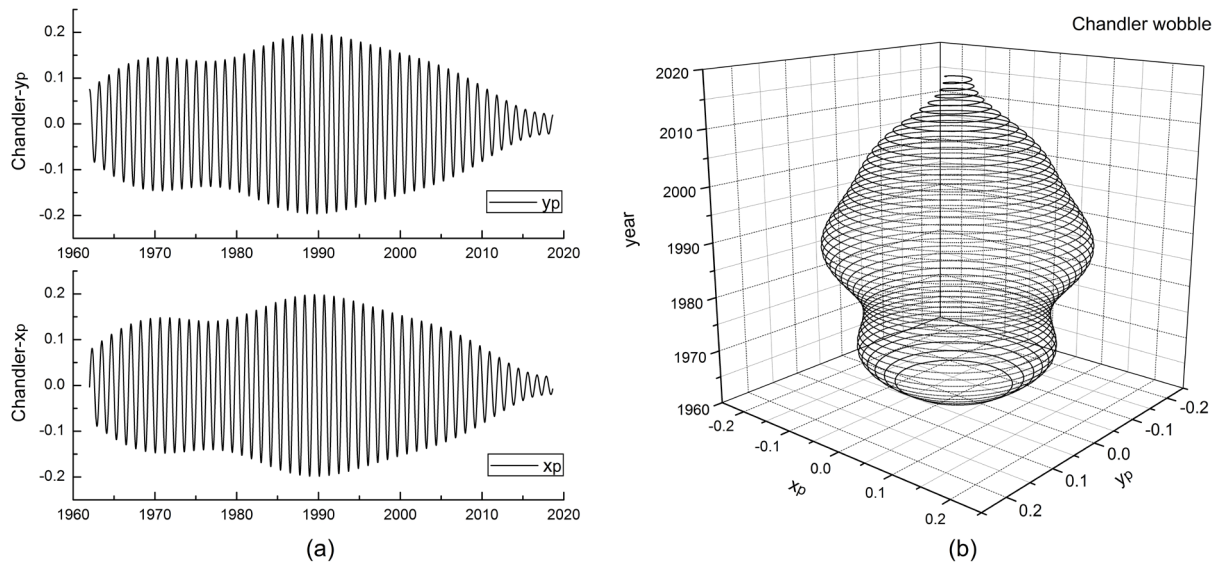


Fig. 3. Chandler wobble since 1962: extracted from EOP C04, (a) Two components xp and yp of Chandler wobble [unit: arcsec], (b) Same time series illustrated as 2-dimensional Chandler wobbling motion in time passage.

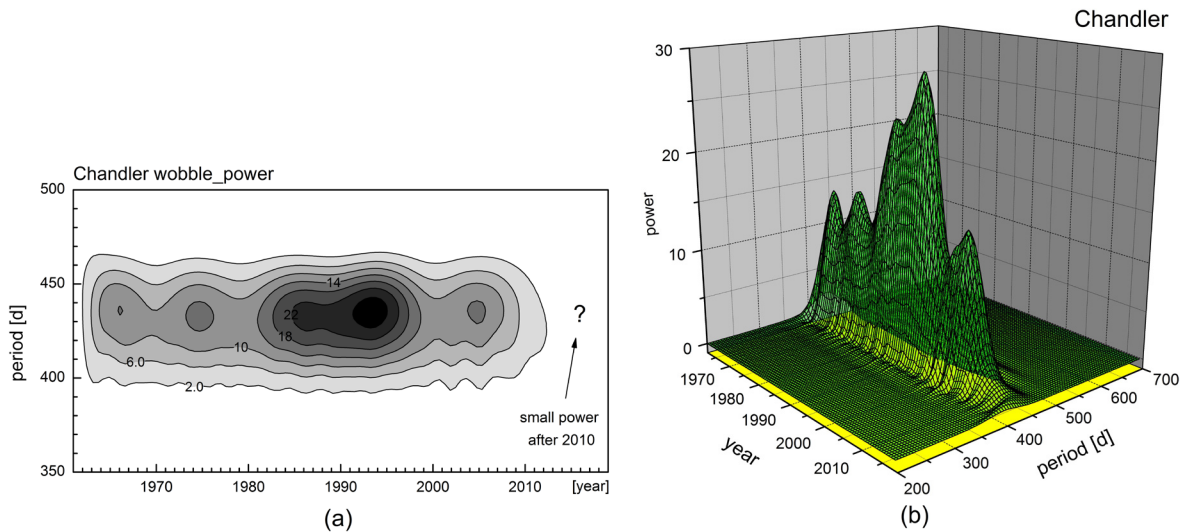


Fig. 4. Morlet wavelet spectrum of Chandler wobble: (a) simple wavelet power spectrum of Chandler wobble with equal-power lines, (b) same spectrum illustrated as three dimensional surface in time passage.

Three different criteria have been tested: (a) minimum sum of squared difference (Gross method; in Plag et al. 2005), (b) minimum sum of power difference (Furuya & Chao method; Furuya & Chao, 1996), (c) minimum sum of power difference plus total power difference. Results are followings: (a) $T = 431.0$ days & $Q = 265$, (b) $T = 432.5$ days & $Q = 47$, (c) two sets $T = 432.3$ days & $Q = 45$ and $T = 433.1$ days & $Q = 43$. The results are also illustrated in Fig. 5. Simple average of these may be taken as $T = 432.2$ days and $Q = 100$.

3.2 FEATURES OF FREE CORE NUTATION

The celestial pole offset; δX and δY since 1984 are given

in EOP C04 dataset. In Fig. 6, the data and the free core nutation as acquired by filtering on the frequency domain are shown together. Relatively large error in δX and δY exist in the early time before 1990. Basically the celestial pole off set has been determined from vast amount of VLBI observations through lengthy reduction procedure including the comparison with model of the Earth's precession and nutation.

The Fourier power spectrum of the celestial pole offset is shown in Fig. 7. For better resolution of the spectrum we again used four sets of day number = $4k$, $4k+1$, $4k+2$, and $4k+3$, and then added the each four corresponding spectra. From the spectra, it is clear that free core nutation takes

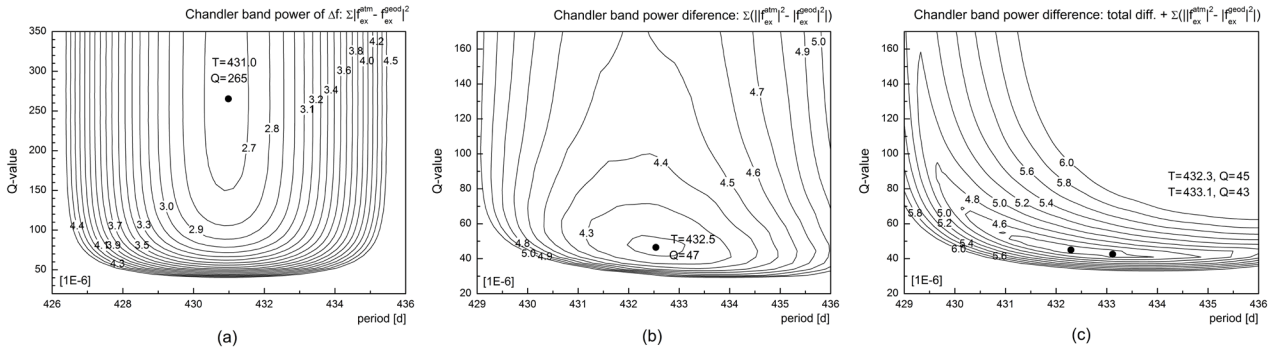


Fig. 5. Estimates of Chandler period and quality factor acquired by least square error of spectrum in Chandler band. Three kinds of criteria used: (a) minimum sum of squared difference (Gross method), (b) minimum sum of power difference (Furuya & Chao method), (c) minimum sum of power difference plus total power difference. Each identified period and Q-value set for minimum square error are $T = 431.0$ days & $Q = 265$, $T = 432.5$ days & $Q = 47$, and two sets $T = 432.3$ days & $Q = 45$ and $T = 433.1$ days & $Q = 43$.

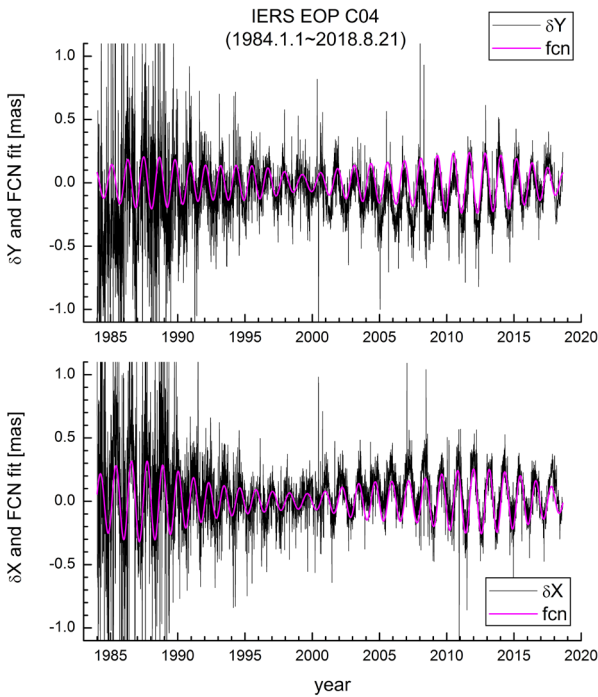


Fig. 6. The celestial pole offset; δX and δY since 1984 determined from VLBI observation. Free core nutation component acquired by filtering is imposed on the data [unit: milliarcsec].

more than 95 percent of total power, however, small content of 18.6 year and 1 year periodicities as well as other weak periodicities can be seen (Fig. 7a). These minor contents can be regarded as minute traces of residual forced nutation - not compensated by the model. Fourier spectrum shows that spectral peaks are split into three ones, of which periods are 448, 423, and 400 days and others (Fig. 7b). Its center period was found as 435.6 days ($409.2d < T < 465.6d$ as $\pm 1\sigma$ range) from the distribution of power in frequency domain. Assuming white spectral excitation, we found the quality factor of free core nutation as $Q = 8$ from the power

spectrum. Corresponding quality factor of nearly diurnal free wobble follows as $Q = 3600$.

Wavelet spectrum of celestial pole offset time series has been attained by using Morlet wavelet and is illustrated in Fig.8. Large variations of both the temporal amplitude and periodicity of free core nutation are noticeable. 3-dimensional illustration gives panoramic view of these features.

4. DISCUSSION AND CONCLUSION

As shown in Fig. 3a the amplitude of Chandler wobble kept on decreasing since 1990 and is now one third of its average amount or less. Though not shown here, atmospheric excitation has not so decreased with such large reduction at the same time, therefore, other source of excitation must have affected significantly. From the comparisons between geodetic excitation and atmospheric excitation, the period and Q-value of Chandler wobble were determined as $T = 432.2$ days and $Q = 100$ in this study, while direct interpretation on the power spectrum led $T = 434.3$ and $Q = 30$. This discrepancy could be partly due to un-modeled oceanic excitation.

Average period of free core nutation was found as $T = 435.6$ days in the given time span (1984 - 2018), however, its periodicity is found unstable with variations over 50 days. Unlike Chandler wobble, we have not attempted to infer the period and Q-value of free core nutation by excitation comparison, for its excitation function is not quite accurately formulated. From the power spectrum the quality factor of free core nutation was read as $Q = 8$, which corresponds to nearly diurnal free wobble quality factor as $Q = 3600$. Morlet wavelet spectrum well describes the amplitude and period variations of free core nutation.

Existence of free core nutation is due to misalignment

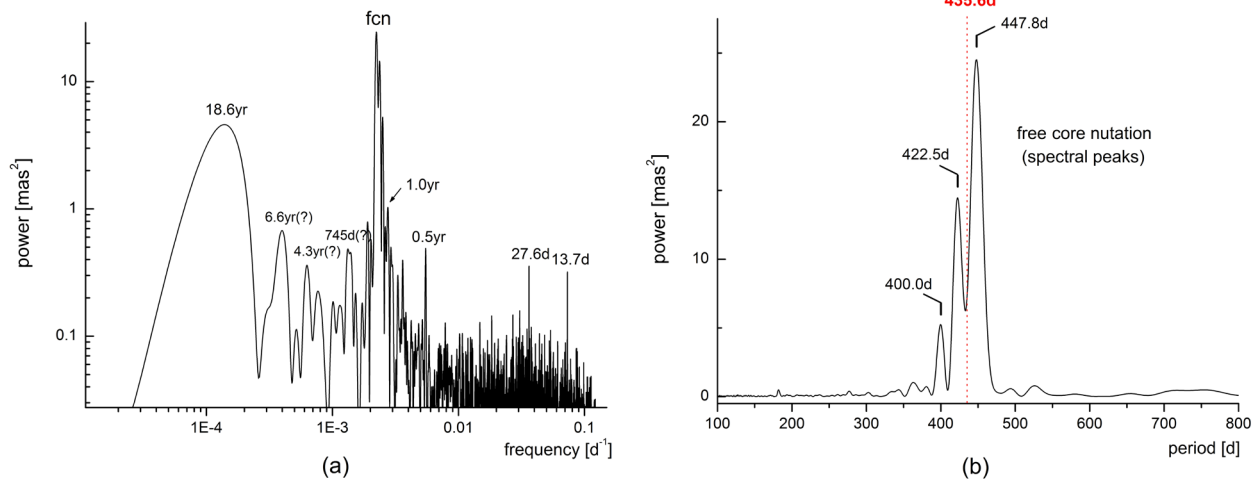


Fig. 7. Fourier power spectra of the pole offset δX and δY shown in Fig. 6: (a) power spectrum shown on frequency axis, (b) spectral peaks of free core nutation shown on period axis. The pole offset data contain small amounts of other components. Free core nutation peaks are split and its average period is 435.6 days.

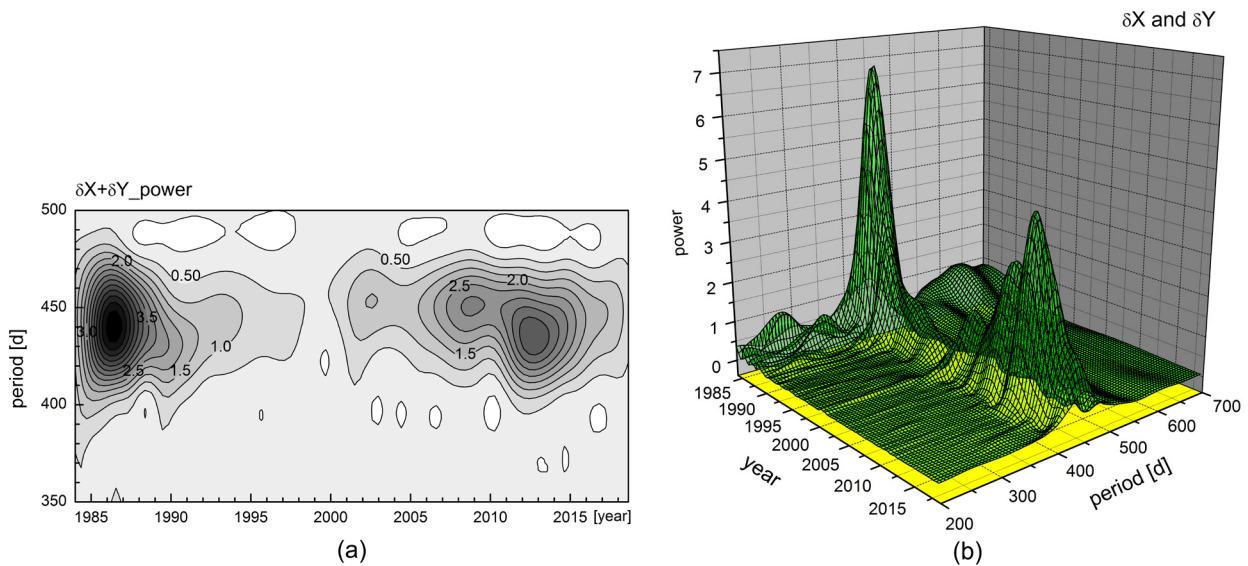


Fig. 8. Morlet wavelet spectrum of the free core nutation time series (celestial pole offset δX and δY): (a) wavelet power spectrum shown with equal-power lines, (b) same spectrum illustrated as three dimensional surface in time passage.

of rotation axes of mantle and liquid core associated with the ellipticity of core-mantle boundary ($\epsilon = 0.00252$ in Eq. 10). Thus its large period variation is quite possibly caused by a certain change in the core-mantle boundary. More investigations are needed in both its theory and data analysis. It is also desirable to have reliable formulation for excitation function of free core nutation.

ACKNOWLEDGMENTS

This study has been financially supported by Korea Astronomy and Space Science Institute.

REFERENCES

Chao BF, Hsieh Y, The Earth's free core nutation: Formulation of dynamics and estimation of eigenperiod from the very-long-baseline interferometry data, *Earth Plan. Sci. Lett* 432, 483-495 (2015). <http://dx.doi.org/10.1016/j.epsl.2015.10.010>

Chung TW, Na SH, A Least Square Fit Analysis on the Earth's Polar Motion Time Series: Implication against Smylie's Conjecture, *Geophys. Geophys. Explor.* 19, 91-96 (2016). <http://dx.doi.org/10.7582/GGE.2016.19.2.091>

- Cummins PR, Wahr JM, A Study of the Earth's free core nutation using deployment of accelerometers gravity data, *J. Geophys. Res.* 98, 2091-2103 (1983). <http://dx.doi.org/10.1029/92JB01956>
- Defraigne P, Dehant V, Hinderer J, Stacking gravity tide measurement and nutation observations in order to determine the complex eigenfrequency of the nearly diurnal free wobble, *J. Geophys. Res.* 99, 9203-9213 (1994). <http://dx.doi.org/10.1029/94JB00133>
- Furuya M, Chao BF, Estimation of period and Q of Chandler wobble, *Geophys. J. Intern.* 127, 693-702 (1996). <https://dx.doi.org/10.1111/j.1365-246X.1996.tb04047.x>
- Gross RS, Earth Rotation Variations - Long Period, in *Treatise of Geophysics*, vol. 3, Geodesy (Elsevier, Amsterdam, 2009), 239-294.
- Gubanov VS, New estimates of retrograde free core nutation parameters, *Astron. Lett.* 36, 444-451 (2010). <http://dx.doi.org/10.1134/S1063773710060083>
- Krasna H, Boehm J, Schuh H, Free core nutation observed by VLBI, *Astronomy & Astrophysics.* 555, A29 (2013). <http://dx.doi.org/10.1051/0004-6361/201321585>
- Mathews PM, Herring TA, Buffett BA, Modelling of nutation and precession: New nutation series for nonrigid Earth and insights into the Earth's interior, *J. Geophys. Res.* 107, TEC3 (2002), <http://dx.doi.org/10.1029/2001JB000390>
- Moritz H, Mueller II, *EARTH ROTATION: Theory and Observation*, (Frederick Ungar, New York, 1988).
- Munk WH, MacDonald GJE, *The Rotation of the Earth: A Geophysical Discussion* (Univ. Cambridge Press, Cambridge, 1960).
- Plag HP, Chao BF, Gross RS, Van Dam T, *Proceedings of the Workshop on Forcing of polar motion in the Chandler frequency band: a contribution to understanding interannual climate variations*, The European Center for Geodynamics and Seismology, Luxembourg, 13-15 April 2005.
- Petit, G, Luzum, B, *IERS Conventions* (2010), IERS Technical Note, No. 36, (2010).
- Rajner M, Brzezinski A, Free core nutation period inferred from the gravity measurements at Jozefoslaw, *Stud. Geophys. Geod.* 61, 639-656 (2017), <http://dx.doi.org/10.1007/s11200-016-0174-4>
- Sasao T, Okubo S, Saito M, A simple theory on dynamical effects of stratified fluid core upon nutational motion of the Earth, in 1977 IAU Symposium, Kiev, USSR, 23-28 May 1977.
- Smith ML, Wobble and nutation of the Earth, *Geophys. J. Intern.* 50, 103-140 (1977). <https://dx.doi.org/10.1111/j.1365-246X.1977.tb01326.x>

APPENDIX: Celestial/Terrestrial Coordinates – Precession, Nutation and Polar Motion

Coordinate transformation from Celestial Reference Frame to Terrestrial Reference Frame has been mostly expressed as follows (IAU 2000A), while the transform given in the introduction is equivalent one using two celestial pole coordinate X and Y for precession plus nutation.

$$\begin{bmatrix} x_1 \\ x_2 \\ x_3 \end{bmatrix}_{\text{Terrestrial}} = R_{pm} R_{spin} R_{nut} R_{prec} \begin{bmatrix} x_1 \\ x_2 \\ x_3 \end{bmatrix}_{\text{Celestial}} \quad (\text{A1})$$

R_{prec} , R_{nut} , R_{spin} , and R_{pm} in Eq.(A1) are each rotation matrices corresponding to the precession, nutation, Earth's spin, and polar motion.

$$\begin{aligned} R_{prec} &= R_3(-z)R_2(\theta)R_3(-\zeta) \\ R_{nut} &= R_1(-\varepsilon - \Delta\varepsilon)R_3(-\Delta\psi)R_1(\varepsilon) \\ R_{spin} &= R_3(GAST) \\ R_{pm} &= R_2(-x_p)R_1(-y_p) \end{aligned}$$

The three angles of the precession matrix; $R_{prec} = R_3(-z)R_2(\theta)R_3(-\zeta)$ are given as follows.

$$\begin{aligned} \zeta &= 2.5976176'' + 2306.0809506''T + 0.3019015''T^2 \\ &\quad + 0.0179663''T^3 - 0.0000327''T^4 - 0.0000002''T^5 \\ \theta &= 2004.1917476''T - 0.4269353''T^2 - 0.0418251''T^3 \\ &\quad - 0.0000601''T^4 - 0.0000001''T^5 \\ z &= -2.5976176'' + 2306.0803226''T + 1.0947790''T^2 \\ &\quad + 0.0182273''T^3 + 0.0000470''T^4 - 0.0000003''T^5 \end{aligned}$$

For precession, one may assume the lunar and solar masses as circularly distributed around the Earth like donuts. In fact, the moon and the sun give periodic torques as oscillatory perturbations and lead Earth nutation. By analogy of harmonic oscillator to periodic forces, amplitude

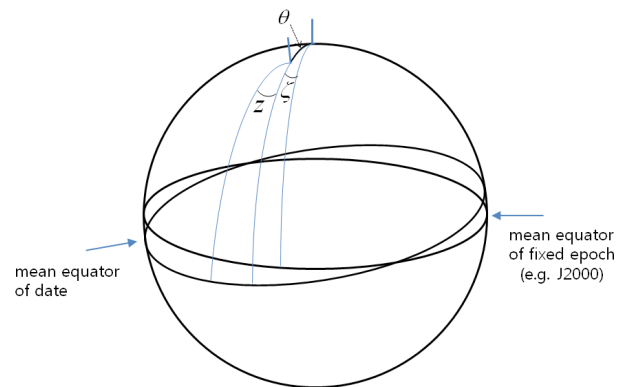


Fig. A1. Earth's precession represented by three angles ζ , θ , and z . Precession matrix is given as $R_3(-z)R_2(\theta)R_3(-\zeta)$.

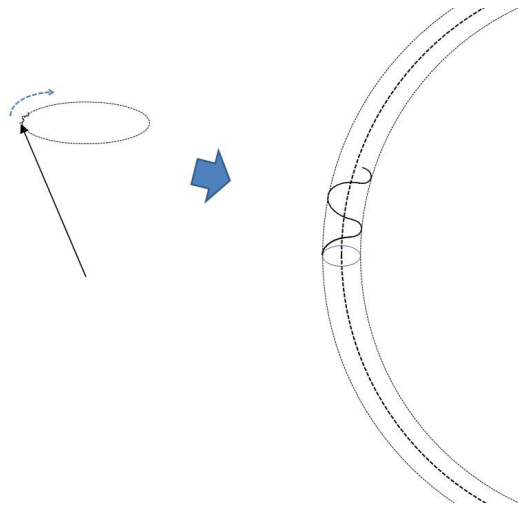


Fig. A2. 18.6 year nutation superposed on the precession. Approximate locus of celestial intermediate pole for about 1.5 nutation period is drawn.

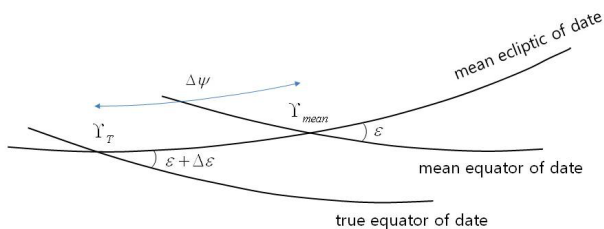


Fig. A3. Nutation angles $\Delta\epsilon$ and $\Delta\psi$.

of long period nutation is larger than short period one.

In Fig. A2, the largest nutation (18.6 year period: the retrograde precession of the lunar orbital plane) is illustrated together with the precession. For different nutation components of same origin (lunar or solar), lower frequency nutation has larger amplitude due to the inverse frequency dependence of nutation angular momentum; $L_n = |\vec{L}_n| = |\tau_n|/\omega_n$. Semiannual nutation is roughly six times larger than fortnightly one (lunar tidal force is about twice larger). Due to symmetric nature of tidal force, semi-annual nutation is of larger amplitude than annual nutation (particularly for $\Delta\epsilon$), and fortnightly nutation amplitude is larger than monthly. If the lunar/solar orbits were completely circular, then the monthly and annual nutation will no longer exist. The two angles $\Delta\epsilon$ and $\Delta\psi$, illustrated in Fig. A3, represent nutation and largest nutation components are listed in the following table. For comparison the amplitudes of four largest nutation are shown in Fig. A4.

Transformation matrix of the nutation described by $\Delta\epsilon$ and $\Delta\psi$ is given as $R_{nut} = R_1(-\epsilon-\Delta\epsilon)R_3(-\Delta\psi)R_1(\epsilon)$. Next to the largest 18.6 year nutation, large ones are semiannual, fortnightly, annual and monthly nutation. Mathews model for the precession and nutation has been adopted by International Astronomical Union (IAU) as its current

Table 1. Amplitudes of six largest nutation components [unit: milliarcsec]

period (day)	nutation in obliquity $\Delta\epsilon$	nutation in longitude $\Delta\psi \sin \epsilon$	remark
6798.4	9203	6858	lunar orbital precession
365.3	5	57	annual
182.6	574	526	semi annual
27.6	1	28	monthly
13.7	98	91	fortnightly
9.1	13	12	modulated

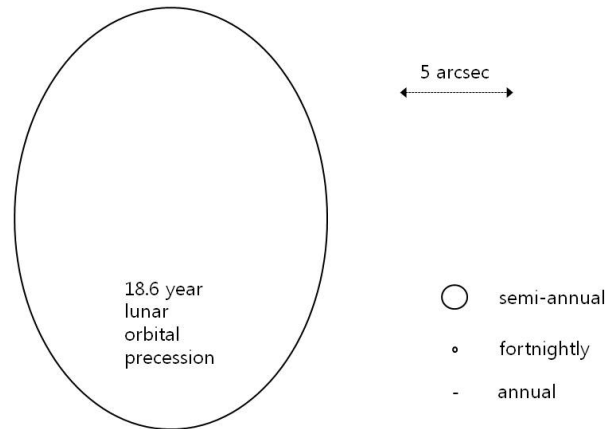


Fig. A4. The nutation ellipses of four major components; 18.6 year, semiannual, fortnightly, and annual. A scale arrow of 5 arcsec is shown.

standard model (Mathews et al. 2002).

Greenwich Apparent Sidereal Time (GAST) is defined as $GAST = GMST + \Delta\psi \cos \epsilon$, where $\Delta\psi \cos \epsilon$ is the nutation of right ascension, while GMST is the Greenwich Mean Sidereal Time in radian defined as $GMST = GMST_0 + \alpha UT1$ with $GMST_0$ given in second as follows.

$$GMST_0 = 24110.54841 + 8640184.812866T + 0.093104T^2 - 6.2 \times 10^{-6}T^3 \text{ (second)}$$

where T is time in Julian century (36525 days) from J2000.0, and α is the sidereal time conversion factor defined as $\alpha = 1.002737909350795 + 5.9006 \times 10^{-11}T - 5.9 \times 10^{-15}T^2$.

Unlike other Earth rotation parameters above, two coordinates (xp, yp) of the instantaneous Earth pole offset cannot be accurately predicted in advance, therefore they are determined through observation. The coordinate transformation from Terrestrial Reference Frame to Celestial Reference Frame can be expressed as inverse of the formula (A1).